2 Current conservation for Schrödinger equation

Schrödinger wave equation (for one particle):

$$\left(-\frac{\hbar^2}{2m}\Delta + V(x)\right) = i\hbar\frac{\partial\psi}{\partial t}$$
(2.1)

$$H\psi(\vec{x},t) = i\hbar \frac{\partial \psi}{\partial t} \tag{2.2}$$

From Eq.(2.1) and its complex conjugate (*), we get

$$\begin{array}{ll} (H\psi) &=& i\hbar\frac{\partial\psi}{\partial t} \Rightarrow \psi^* \left(H\psi\right) = i\hbar\psi^*\frac{\partial\psi}{\partial t} \\ (H\psi^*) &=& -i\hbar\frac{\partial\psi^*}{\partial t} \Rightarrow \psi \left(H\psi^*\right) = -i\hbar\psi\frac{\partial\psi^*}{\partial t} \end{array}$$

Taking the difference of these two equations, we get

$$\psi^* (H\psi) - \psi (H\psi^*) = i\hbar \frac{\partial}{\partial t} (\psi^* \psi)$$
$$-\frac{\hbar^2}{2m} (\psi^* \Delta \psi - \psi \Delta \psi^*) = i\hbar \frac{\partial}{\partial t} (\psi^* \psi)$$
$$-\frac{\hbar^2}{2m} \vec{\nabla} \cdot \left(\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^*\right) = i\hbar \frac{\partial}{\partial t} (\psi^* \psi)$$

This gives the "current conservation" (continuity equation):

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0 \tag{2.3}$$

where ρ is the probability density

$$\rho = \psi^* \, \psi$$

and \vec{j} is the probability current

$$\vec{j} = \frac{\hbar}{2mi} \left(\psi^* \vec{\nabla} \psi - \psi \vec{\nabla} \psi^* \right)$$

From Eq.(2.3) we get the "conservation of probability"

$$\frac{\partial}{\partial t} \int \mathrm{d}^3 x \, \rho(\vec{x}, t) = 0$$

Important point: Schrödinger equation (2.1) is of first order in time!

<u>Plane wave solutions</u> of (2.1) for free particles:

$$\psi(\vec{x},t) = e^{-i(Et - \vec{p} \cdot \vec{x})/\hbar} \tag{2.4}$$

where \vec{p} are the eigenvalues of the momentum operator $-i\hbar \vec{\nabla}$, and $E = \vec{p}^2/(2m)$ is the energy.

3 Dirac equation

Wave equation of Schrödinger form (first order in $\partial/\partial t$):

$$H\psi(\vec{x},t) = i\hbar \frac{\partial\psi}{\partial t} \tag{3.1}$$

What is the relativistic Hamiltonian H for a free particle ? - In theory of relativity, space and time enter very symmetrically (see, for example, Lorentz transformation).

 \Rightarrow H should be of first order in $\vec{\nabla}$ (or: first order in momentum operator $\vec{p} = -i\hbar\vec{\nabla}$), and have dimension of energy. Try

$$H = (\alpha_1 \, p_1) \, c + (\alpha_2 \, p_2) \, c + (\alpha_3 \, p_3) \, c + \beta \, (mc^2) \equiv (\vec{\alpha} \cdot \vec{p}) \, c + \beta \, (mc^2) \tag{3.2}$$

where $\vec{\alpha} = (\alpha_1, \alpha_2, \alpha_3)$ and β are dimensionless constants (independent of \vec{x} and t.)

<u>But:</u> If $\alpha_1, \alpha_2, \alpha_3; \beta$ are just numbers, then H is not invariant under rotations !

Later, we will see that this problem can be solved if $\alpha_1, \alpha_2, \alpha_3; \beta$ are not constant numbers, but constant <u>matrices</u> ! In this case, ψ is a <u>vector</u>!

Take $i\hbar(\partial/\partial t)$ of wave equation (3.1):

$$-\hbar^{2} \frac{\partial^{2} \psi}{\partial t^{2}} = \left(\left(\vec{\alpha} \cdot \vec{p} \right) c + \beta \left(mc^{2} \right) \right)^{2} \psi$$

$$= \left[c^{2} \left(\vec{\alpha} \cdot \vec{p} \right) \left(\vec{\alpha} \cdot \vec{p} \right) + mc^{3} \left(\left(\vec{\alpha} \cdot \vec{p} \right) \beta + \beta \left(\vec{\alpha} \cdot \vec{p} \right) \right) + \left(mc^{2} \right)^{2} \beta^{2} \right] \psi$$
(3.3)

Here $\vec{p} = -i\hbar\vec{\nabla}$ is the momentum operator.

The solution for a free particle (momentum \vec{p}) must be of the form

$$\psi(\vec{x},t) = w(\vec{p}) e^{-i(Et - \vec{p} \cdot \vec{x})/\hbar}$$
(3.4)

where $w(\vec{p})$ ist a vector, independent of (\vec{x}, t) .

Inserting (3.4) into (3.3) we get

$$E^{2} = c^{2} \left(\vec{\alpha} \cdot \vec{p} \right) \left(\vec{\alpha} \cdot \vec{p} \right) + mc^{3} \left(\left(\vec{\alpha} \cdot \vec{p} \right) \beta + \beta \left(\vec{\alpha} \cdot \vec{p} \right) \right) + (mc^{2})^{2} \beta^{2} \right)$$
(3.5)

This must be the energy squared of a free particle with momentum \vec{p} , i.e., $E^2 = (pc)^2 + (mc^2)^2$. Therefore

$$\left(\vec{\alpha} \cdot \vec{p}\right) \left(\vec{\alpha} \cdot \vec{p}\right) = \vec{p}^2 \tag{3.6}$$

$$\left(\vec{\alpha} \cdot \vec{p}\right)\beta + \beta\left(\vec{\alpha} \cdot \vec{p}\right) = 0 \tag{3.7}$$

$$\beta^2 = 1 \tag{3.8}$$

This must hold for any vector \vec{p} , therefore

$$\alpha_i \alpha_j + \alpha_j \alpha_i = 2\delta_{ij} \qquad (i, j = 1, 2, 3) \tag{3.9}$$

$$\alpha_i \beta + \beta \alpha_i = 0 \quad (i = 1, 2, 3)$$
(3.10)

$$\alpha_i^2 = 1$$
 $(i = 1, 2, 3), \quad \beta^2 = 1$ (3.11)

Conditions for matrices $\alpha_1, \alpha_2, \alpha_3, \beta$:

- Hermite (because Hamiltonian *H* must be hermite).
- Eigenvalues are ± 1 (because of Eq.(3.11)).
- Trace is zero: For example, from Eq.(3.10), $\operatorname{Tr}(\alpha_i) = -\operatorname{Tr}(\beta \alpha_i \beta) = -\operatorname{Tr}(\beta^2 \alpha_i) = -\operatorname{Tr}(\alpha_i) = 0.$

Therefore, sum of eigenvalues must be zero \Rightarrow dimension of matrices must be even (N = 2, 4, 6, ...). But N = 2 is not possible: The four independent hermite 2×2 matrices are the Pauli matrices $\vec{\sigma} = (\sigma_1, \sigma_2, \sigma_3)$ and the unit matrix 1. But then (3.10) is not satisfied.

Lowest dimension is N = 4. For example, the following matrices satisfy all conditions (3.9) - (3.11):

$$\vec{\alpha} = \begin{pmatrix} 0 & \vec{\sigma} \\ \vec{\sigma} & 0 \end{pmatrix}, \qquad \beta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}$$
(3.12)

<u>Home work</u>: Show that these matrices satisfy the conditions (3.9) - (3.11).

Current conservation for Dirac equation:

Dirac equation (3.1) and its hermite conjugate (\dagger) :

$$i\hbar\frac{\partial\psi}{\partial t} = -i\hbar c\,\vec{\alpha}\cdot\left(\vec{\nabla}\psi\right) + \beta\,mc^2\,\psi \tag{3.13}$$

$$-i\hbar \frac{\partial \psi^{\dagger}}{\partial t} = i\hbar c \left(\vec{\nabla}\psi^{\dagger}\right) \cdot \vec{\alpha} + \psi^{\dagger} \beta mc^{2}$$
(3.14)

Multiply (3.13) from left by ψ^{\dagger} , and (3.14) from right by ψ , and take the difference of these equations:

$$i\hbar\frac{\partial}{\partial t}\left(\psi^{\dagger}\psi\right) = -i\hbar c\,\vec{\nabla}\cdot\left(\psi^{\dagger}\vec{\alpha}\psi\right)$$

This has the form of <u>current conservation</u>:

$$\frac{\partial \rho}{\partial t} + \vec{\nabla} \cdot \vec{j} = 0 \tag{3.15}$$

with the probability density

$$\rho = \psi^{\dagger}\psi \tag{3.16}$$

and the probability current

$$\vec{j} = c\psi^{\dagger}\vec{\alpha}\psi \tag{3.17}$$

Relation to the spin:

- There must be <u>four</u> independent solutions to the Dirac equation (3.1).
- The are <u>two</u> different energy eigenvalues: From (3.5), $E = \pm \sqrt{(pc)^2 + (mc^2)^2}$.
- Therefore, two solutions have energy $E = +\sqrt{(pc)^2 + (mc^2)^2}$, and two solutions have energy $E = -\sqrt{(pc)^2 + (mc^2)^2}$.

This <u>2-fold degeneracy</u> means <u>spin 1/2</u>: For a free particle, spin up (\uparrow) and spin down (\downarrow) have the same energy.

The angular momentum operator for the Dirac equation is given by

$$\vec{J} = \vec{L} + \vec{S} = \vec{r} \times \vec{p} + \frac{\hbar}{2}\vec{\Sigma}$$
(3.18)

where the 4×4 matrices $\vec{\Sigma} = (\Sigma_1, \Sigma_2, \Sigma_3)$ are given by

$$\vec{\Sigma} = \begin{pmatrix} \vec{\sigma} & 0\\ 0 & \vec{\sigma} \end{pmatrix} \tag{3.19}$$

The commutation relations of \vec{L} and \vec{S} with the Hamiltonian $H = c(\vec{\alpha} \cdot \vec{p}) + \beta mc^2$ are as follows:

$$[H, L_i] = -i\hbar c \left(\vec{\alpha} \times \vec{p}\right)_i$$
$$[H, S_i] = i\hbar c \left(\vec{\alpha} \times \vec{p}\right)_i$$

Try to check these relations!

Therefore $\left[H, \vec{J}\right] = 0$, and the <u>total</u> angular momentum \vec{J} (sum of orbital and spin angular momentum) is conserved !